Atoms and photons Chapter 4

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September 12, 2016

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2 Spontaneous emission in free space

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2 Spontaneous emission in free space



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#### Outline



2 Spontaneous emission in free space

#### 3 Photodetection



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#### Interaction of quantum light with matter

- Quantum field and classical charges
- Quantum field and quantized atom

Coupling of a quantum mode with a classical current (model of electronic source)

$$\mathbf{j}(\mathbf{r},t) = \mathbf{j}_0(\mathbf{r})e^{-i\omega_0 t} \tag{1}$$

Simplified field Hamiltonian

$$H_0' = \sum_{\ell} \hbar \omega_{\ell} a_{\ell}^{\dagger} a_{\ell}$$
 (2)

From classical interaction energy,  $-\mathbf{j}\cdot\mathbf{A}$ , guess the interaction Hamiltonian

$$H_i = -\int_{\mathcal{V}} \mathbf{j}(\mathbf{r}, t) \cdot \mathbf{A}(\mathbf{r}, t) d^3 \mathbf{r}$$
(3)

where

$$\mathbf{A}(\mathbf{r},t) = \sum_{\ell} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\ell} \mathcal{V}}} a_{\ell} \mathbf{f}_{\ell}(\mathbf{r}) + \text{c.c.}$$
(4)

Interaction representation

$$\left| \widetilde{\Psi} \right\rangle = U_0^{\dagger} \left| \Psi \right\rangle$$
 (5)

with

$$U_0 = e^{-iH'_0 t/\hbar} = \prod_{\ell} e^{-i\omega_{\ell} t a_{\ell}^{\dagger} a_{\ell}}$$
(6)

New Hamiltonian

$$\widetilde{H} = U_0^{\dagger} H_i U_0 \tag{7}$$

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Annihilation operator transformation

$$\widetilde{a}_{\ell} = e^{i\omega_{\ell}ta^{\dagger}_{\ell}a_{\ell}}a_{\ell}e^{-i\omega_{\ell}ta^{\dagger}_{\ell}a_{\ell}}$$

$$= a_{\ell} - i\omega_{\ell}ta_{\ell} + \frac{(i\omega_{\ell}t)^{2}}{2}a_{\ell} + \dots$$

$$= a_{\ell}e^{-i\omega_{\ell}t}$$
(8)

using Baker Hausdorff and  $[N_\ell, a_\ell] = -a_\ell$ .

Interaction representation

$$\widetilde{H} = -\int d^{3}\mathbf{r} \left[\mathbf{j}_{0}(\mathbf{r})e^{-i\omega_{0}t} + \mathbf{j}_{0}^{*}(\mathbf{r})e^{i\omega_{0}t}\right]$$
$$\cdot \left[\sum_{\ell} \sqrt{\frac{\hbar}{2\epsilon_{0}\omega_{\ell}\mathcal{V}}} \left(a_{\ell}e^{-i\omega_{\ell}t}\mathbf{f}_{\ell}(\mathbf{r}) + a_{\ell}^{\dagger}e^{i\omega_{\ell}t}\mathbf{f}_{\ell}^{*}(\mathbf{r})\right)\right] \qquad (9)$$

Rotating wave approximation for  $\omega_{\ell} \approx \omega_0$ :

$$\widetilde{H} = -\sum_{\ell} \sqrt{\frac{\hbar \mathcal{V}}{2\epsilon_0 \omega_{\ell}}} J_0 a_{\ell}^{\dagger} e^{-i(\omega_0 - \omega_{\ell})t} + \text{h.c.}$$
(10)

where the complex scalar  $J_0$  is defined as:

$$J_0 = \frac{1}{\mathcal{V}} \int d^3 \mathbf{r} \, \mathbf{j}_0(\mathbf{r}) \cdot \mathbf{f}_\ell^*(\mathbf{r}) \tag{11}$$

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Single mode evolution

Setting

$$K_0 = \sqrt{\frac{\hbar \mathcal{V}}{2\epsilon_0 \omega}} J_0 \tag{12}$$

we can write the Hamiltonian in the simpler form

$$\widetilde{H} = -K_0 e^{-i\delta t} a^{\dagger} + \text{h.c.}$$
(13)

where

$$\delta = \omega_0 - \omega \tag{14}$$

Note that the Hamiltonians at different times do not commute. Evolution operator not simple.

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Single mode evolution

From *t* to t + dt:

$$\widetilde{H} = -K_0 e^{-i\Phi} a^{\dagger} + \text{h.c.}$$
(15)

where

$$\Phi = \delta t \tag{16}$$

The evolution operator is then a displacement:

$$U(t, t + dt) = D(d\alpha)$$
(17)

with

$$d\alpha = \frac{iK_0}{\hbar} e^{-i\Phi} dt \tag{18}$$

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#### Quantum field and classical charges Single mode evolution

By adding up the amplitudes and within a global phase, the final state is a coherent state with amplitude

$$\beta = \int_0^t \frac{iK_0}{\hbar} e^{-i\delta t'} dt' = -\frac{K_0}{\hbar\delta} \left[ e^{-i\delta t} - 1 \right]$$
(19)

- For  $\delta \neq 0$ , periodic variation of the amplitude
- For  $\delta = 0$  $\beta = \frac{iK_0}{\hbar}t$

Linear amplitude and quadratic photon number growth

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# Quantum field and quantized atom

Hamiltonians

$$H_{ap} = -\frac{q}{m} \mathbf{P} \cdot \mathbf{A}(0)$$
(21)  
$$H_{de} = -\mathbf{D} \cdot \mathbf{E}(0)$$
(22)

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## Quantum field and quantized atom

Electric dipole interaction

Dipole

$$\mathbf{D} = d\epsilon_d |g\rangle \langle e| + \text{h.c.}$$
(23)

Electric field in the plane mode basis

$$\mathsf{E}(0) = i \sum_{\ell} \sqrt{\frac{\hbar \omega_{\ell}}{2\epsilon_0 \mathcal{V}}} a_{\ell} \epsilon_{\ell} + \text{h.c.}$$
(24)

For nearly resonant modes (dominant effect in general), two of the four terms in  $\mathbf{D} \cdot \mathbf{E}(0)$  can be neglected (RWA approximation)

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#### Spontaneous emission

Coupling an atom to the continuum of modes in free space. Decay of the excited states and (diverging) shifts of the energy levels.

- Fermi Golden Rule argument
- Wigner Weisskopf calculation

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#### Fermi Golden rule

Initial state  $|e, 0\rangle$ . Continuum of final states  $|g, 1_{\ell}\rangle$ . Compute separately the rate of photon emission in all directions:

$$\Gamma = \int d\Gamma d\Omega \tag{25}$$

$$d\Gamma = \sum_{\boldsymbol{\epsilon}_{\ell}} \frac{2\pi}{\hbar} |W|^2 d\rho (\boldsymbol{E} = \hbar\omega_0, d\Omega)$$
(26)

Density of states  $d\rho=\rho d\Omega/4\pi$  where

$$\rho(\nu) = \frac{8\pi}{2c^3} \mathcal{V}\nu^2 d\nu \tag{27}$$

With  $\rho(E)dE = \rho(\nu)d\nu$  for  $E = h\nu = \hbar\omega_0$ 

$$\rho(E) = \frac{\mathcal{V}}{2\pi^2 c^3} \frac{1}{\hbar} \left(\frac{E}{\hbar}\right)^2 \tag{28}$$

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## Fermi Golden rule

Finally

$$d\rho(E = \hbar\omega_0, d\Omega) = \frac{\mathcal{V}}{8\pi^3} \frac{\omega_0^2}{\hbar c^3} d\Omega$$
<sup>(29)</sup>

Coupling

$$|W|^{2} = |\langle g, 1_{\ell} | \mathbf{D} \cdot \mathbf{E} | e, 0 \rangle|^{2}$$
(30)

Without loss of generality

$$\boldsymbol{\epsilon}_d = \mathbf{u}_z \tag{31}$$

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$$W|^{2} = \left| d\mathbf{u}_{z} \cdot \boldsymbol{\epsilon}_{\ell}^{*} \sqrt{\frac{\hbar\omega_{\ell}}{2\epsilon_{0}\mathcal{V}}} \right|^{2}$$
(32)

We can now evaluate the rate

$$d\Gamma = \sum_{\boldsymbol{\epsilon}_{\ell}} \frac{1}{8\pi^2 \epsilon_0} \frac{\omega_0^3}{c^3} \frac{|\boldsymbol{d}|^2}{\hbar} |\mathbf{u}_{\boldsymbol{z}} \cdot \boldsymbol{\epsilon}_{\ell}^*|^2 d\Omega$$
(33)

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#### Fermi Golden rule

Expand  $\mathbf{u}_z$  on the basis of  $\mathbf{u}_k$  (propagation direction) and two orthogonal linear polarizations  $\epsilon_1$  and  $\epsilon_2$ :

$$(\mathbf{u}_z \cdot \boldsymbol{\epsilon}_1^*)^2 + (\mathbf{u}_z \cdot \boldsymbol{\epsilon}_2^*)^2 = 1 - (\mathbf{u}_z \cdot \mathbf{u}_k)^2 = 1 - \cos^2 \theta = \sin^2 \theta \qquad (34)$$

Integration over solid angle:

$$\Gamma = \frac{1}{8\pi^2\epsilon_0} \frac{\omega_0^3}{c^3} \frac{|d|^2}{\hbar} \int_0^{2\pi} \int_0^{\pi} \sin^3\theta \, d\theta d\phi$$
(35)

and finally

$$\overline{\phantom{a}} = \frac{\omega_0^3 |d|^2}{3\pi\omega_0\hbar c^3} \tag{36}$$

Already used many times in these lectures!

A more detailed insight. Atom-field state at time t:

$$|\Psi(t)\rangle = c_0(t) |e,0\rangle + \sum_{\ell} c_{\ell}(t) |g,1_{\ell}\rangle$$
(37)

Schrödinger equation:

$$i\hbar \frac{dc_0}{dt} = \hbar \omega_0 c_0 + \sum_{\ell} V_{\ell} c_{\ell}$$
(38)  
$$i\hbar \frac{dc_{\ell}}{dt} = \hbar \omega_{\ell} c_{\ell} + V_{\ell}^* c_0$$
(39)

with

$$V_{\ell} = -\langle e, 0 | \mathbf{D} \cdot \mathbf{E} | g, 1_{\ell} \rangle$$
(40)

Interaction representation:

$$b_{\ell} = c_{\ell} e^{i\omega_{\ell} t} \tag{41}$$

$$i\hbar \frac{db_{\ell}}{dt} = e^{i\omega_{\ell}t} V_{\ell}^* c_0 \tag{42}$$

#### Formal integration

$$b_{\ell}(t) = \frac{V_{\ell}^*}{i\hbar} \int_0^t c_0(t') e^{i\omega_{\ell}t'} dt'$$
(43)

#### or

$$c_{\ell}(t) = \frac{V_{\ell}^{*}}{i\hbar} \int_{0}^{t} c_{0}(t') e^{i\omega_{\ell}(t'-t)} dt'$$
(44)

Setting

$$c_0 = e^{-i\omega_0 t} \alpha_0(t) \tag{45}$$

We get

$$\frac{d\alpha_0}{dt} = -\sum_{\ell} \frac{|V_{\ell}|^2}{\hbar^2} e^{i\omega_0 t} \int_0^t e^{i\omega_{\ell}(t'-t)} e^{-i\omega_0 t'} \alpha_0 \, dt' \tag{46}$$

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Changing for the variable au = t - t', we get

$$\frac{d\alpha_0}{dt} = -\int_0^t \mathcal{N}(\tau)\alpha_0(t-\tau)\,d\tau \tag{47}$$

where the integral kernel  ${\cal N}$  is:

$$\mathcal{N}(\tau) = \frac{1}{\hbar^2} \sum_{\ell} |V_{\ell}|^2 e^{i(\omega_0 - \omega_{\ell})\tau}$$
(48)

$$\mathcal{N}(\tau) = \frac{|d|^2}{\hbar^2} \left[ \sum_{\ell} |\mathbf{u}_z \cdot \boldsymbol{\epsilon}_{\ell}|^2 \frac{\hbar \omega_{\ell}}{2\epsilon_0 \mathcal{V}} e^{-i\omega_{\ell} \tau} \right] e^{i\omega_0 \tau}$$
(49)

In a time of the order of  $1/\omega_0$ ,  ${\cal N}$  practically vanishes

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Thus:

$$\int_{0}^{t} \mathcal{N}(\tau)\alpha(t-\tau) d\tau \approx \alpha_{0}(t) \int_{0}^{\infty} \mathcal{N}(\tau) d\tau = \left(\frac{\Gamma}{2} + i\Delta\right) \alpha_{0}(t) \quad (50)$$
$$\frac{d\alpha_{0}}{dt} = -\left(\frac{\Gamma}{2} + i\Delta\right) \alpha_{0} \quad (51)$$

- $\Gamma$  spontaneous emission rate
- $\Delta$  level shift

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#### Final solution

$$c_0(t) = e^{-\Gamma t/2} e^{-i\omega_0 t} e^{-i\Delta t}$$
(52)

$$c_{\ell}(t) = \frac{V_{\ell}}{i\hbar} \frac{1 - e^{-\Gamma t/2} e^{i(\omega_{\ell} - \omega_0 - \Delta)t}}{(\Gamma/2) - i(\omega_{\ell} - \omega_0 - \Delta)}$$
(53)

$$|c_{\ell}(\infty)|^{2} = \frac{|V_{\ell}|^{2}}{\hbar^{2}} \frac{1}{(\Gamma^{2}/4) + (\omega_{\ell} - \omega_{0} - \Delta)^{2}}$$
(54)

a lorentzian profile for the spontaneous emission line.

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Explicit integration of the kernel:

$$\left(\frac{\Gamma}{2} + i\Delta\right) = \frac{|d|^2}{\hbar^2} \sum_{\ell} \left(\mathbf{u}_z \cdot \boldsymbol{\epsilon}_\ell\right)^2 \frac{\hbar\omega_\ell}{2\epsilon_0 \mathcal{V}} \int_0^\infty e^{i(\omega_0 - \omega_\ell)\tau} \, d\tau \qquad (55)$$

Using

$$\int_{0}^{\infty} e^{i\omega t} dt = \pi \delta(t) + i\mathcal{P}\mathcal{P}\frac{1}{\omega}$$
(56)

get for the real part:

$$\Gamma = \frac{2\pi |d|^2}{\hbar^2} \sum_{\ell} \left( \mathbf{u}_z \cdot \boldsymbol{\epsilon}_{\ell} \right)^2 \frac{\hbar \omega_{\ell}}{2\epsilon_0 \mathcal{V}} \delta(\omega_0 - \omega_{\ell})$$
(57)

same result as the Fermi Golden Rule.

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Wigner-Weisskopf
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Level shift

A severe problem

 $\Delta$  is divergent

A (not so simple) solution

Renormalization...

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#### Photodetector model

A simple single system photodetector. A ground state  $|g\rangle$  and a continuum of excited states  $|e_i\rangle$ . Transition to excited state is a click. Detector Hamiltonian

$$\mathcal{H}_{d} = \sum_{i} \hbar \omega_{i} \left| e_{i} \right\rangle \left\langle e_{i} \right| \tag{58}$$

Detector-field interaction  $-\mathbf{D} \cdot \mathbf{E}$  with

$$\mathbf{D} = \sum_{i} d_{i}(\epsilon_{i} |g\rangle \langle e_{i}| + \epsilon_{i}^{*} |e_{i}\rangle \langle g|)$$
(59)

Hence, within irrelevant factors

$$H_{i} = \sum_{i} \kappa_{i} |e_{i}\rangle \langle g| E^{+} + \text{h.c.}$$
(60)

#### Photodetector model

Interaction representation  $a_\ell 
ightarrow a_\ell \exp(-i\omega_\ell t)$ ,  $\ket{e_i}ra{g}
ightarrow \exp(i\omega_i t)\ket{e_i}ra{g}$ 

$$\widetilde{H}_{i} = \sum_{i} \kappa_{i} e^{i\omega_{i}t} |e_{i}\rangle \langle g| E^{+}(t) + \text{h.c.}$$
(61)

Initial condition

$$|\Psi(0)\rangle = |g\rangle \otimes |\Psi_f\rangle$$
 (62)

Image: Image:

State at time t

$$|\Psi(t)\rangle = |g\rangle \otimes |\Psi_f\rangle + \frac{1}{i\hbar} \int_0^t \widetilde{H}_i(t') |\Psi(t')\rangle dt'$$
 (63)

First-order perturbative solution by replacing in the r.h.s.  $|\Psi(t')\rangle$  by  $|\Psi(0)\rangle = |g\rangle \otimes |\Psi_f\rangle$ .

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#### Photodetector model

Noting that, in  $\widetilde{H}_i$ ,  $\ket{g}ig\langle e_i \ket{E^-}$  gives zero on the initial state

$$|\Psi(t)\rangle = |g\rangle \otimes |\Psi_f\rangle + \frac{1}{i\hbar} \sum_{i} \kappa_i \left[ \int_0^t dt' \, e^{i\omega_i t'} E^+(t') \, |\Psi_f\rangle \right] \otimes |e_i\rangle \quad (64)$$

Probability for having a count at time t

$$p_{e} = \sum_{i} |\langle e_{i} | \Psi \rangle|^{2} = \sum_{i} \langle \Psi | e_{i} \rangle \langle e_{i} | \Psi \rangle$$

$$p_{e} = \frac{1}{\hbar^{2}} \sum_{i} |\kappa_{i}|^{2} \int_{0}^{t} dt' \int_{0}^{t'} dt'' e^{i\omega_{i}(t'-t'')} \langle \Psi_{f} | E^{-}(t'')E^{+}(t') | \Psi_{f} \rangle$$

$$(66)$$

#### Photodetector model

For a high density of final states

$$\sum_{i} \longrightarrow \int d\omega \rho(\omega) \tag{67}$$

$$\int d\omega e^{i\omega(t'-t'')} = \pi \delta(t'-t'')$$
(68)

Hence

$$p_e(t) \propto \int_0^t dt' \langle \Psi_f | E^-(t') E^+(t') | \Psi_f \rangle$$
(69)

With a large set of photo-detecting systems the 'photocurrent' is proportional to

$$I(t) = \langle \Psi_f | E^{-}(t) E^{+}(t) | \Psi_f \rangle$$
(70)

Intensity correlations Classical Hanbury-Brown and Twiss

A source, a balanced beamsplitter and two detectors. Correlate the photocurrents:

$$G_2(\tau) = \overline{I_1(t)I_2(t+\tau)} \tag{71}$$

At long times, no correlation.

$$G_2(\infty) = (\overline{I})^2 \tag{72}$$

At  $\tau = 0$ 

$$G_2(0) = \overline{I^2} \tag{73}$$

$$\overline{I^2} - (\overline{I})^2 = \overline{(I - \overline{I})^2} \ge 0$$
(74)

$$G_2(0) \ge G_2(\infty) \tag{75}$$

HBT stellar interferometry

Nature paper (178, 1046): determination of the angular diameter of stars.



HBT stellar interferometry

Measure the intensity correlations at zero delay as a function of the distance d. Envelope field received by  $D_1$ :

$$E_1 = \alpha e^{ikr_{1A}} e^{i\phi_A} + \beta e^{ikr_{1B}} e^{i\phi_B}$$
(76)

with

$$\overline{I_1} = \overline{|E_1|^2} = |\alpha|^2 + |\beta|^2 \tag{77}$$

For  $D_2$ 

$$E_2 = \alpha e^{ikr_{2A}} e^{i\phi_A} + \beta e^{ikr_{2B}} e^{i\phi_B}$$
(78)

After a painful computation and elimination (on the average) of random phases

$$\overline{I_1(t)I_2(t)} = \overline{I_1} \ \overline{I_2} + |\alpha|^2 |\beta|^2 \cos k \left[ r_{1A} + r_{2B} - r_{1B} - r_{2A} \right]$$
(79)

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#### Intensity correlations HBT stellar interferometry

For a symmetric configuration, when A and B are symmetric with respect to the mediating segment  $D_1 - D_2$ ,  $r_{1A} - r_{2A} = r_{2B} - r_{1B}$  and

$$\overline{I_1(t)I_2(t)} = \overline{I_1} \ \overline{I_2} + |\alpha|^2 |\beta|^2 \cos 2k(r_{1A} - r_{2A}) = \overline{I_1} \ \overline{I_2} + |\alpha|^2 |\beta|^2 \cos 2\frac{\omega}{c} (r_{1A} - r_{2A})$$
(80)

or

$$\overline{I_1(t)I_2(t)} = \overline{I_1} \ \overline{I_2} + |\alpha|^2 |\beta|^2 \cos kd\Theta$$
(81)

where  $\Theta$  is the star's angular diameter. Resolution only limited by the distance between the two detectors.

Quantum intensity correlations

#### Admit

$$G_{2}(\mathbf{r}_{1},\mathbf{r}_{2},t,\tau) = \langle \Psi_{f} | \hat{G}_{2} | \Psi_{f} \rangle$$
(82)

where

$$\hat{G}_2 = E^{-}(\mathbf{r}_1, t)E^{-}(\mathbf{r}_2, t+\tau)E^{+}(\mathbf{r}_2, t+\tau)E^{+}(\mathbf{r}_1, t)$$
(83)

Normalized correlation function

$$g_2 = \frac{G_2}{\overline{l_1} \ \overline{l_2}} \tag{84}$$

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Two field modes

Simple situation: two field modes and two detectors.

$$E_i^+ = \mathcal{E}_i e^{i(\mathbf{k}_1 \cdot \mathbf{r} - \omega_i t)} a_i \tag{85}$$

where i = 1, 2.

$$E^+ = E_1^+ + E_2^+ \tag{86}$$

Field state: product of Fock states  $|\Psi
angle=|\textit{N}_1,\textit{N}_2
angle$ . Complex calculation:

- No interference in the simple photocurrents
- Interferences in the  $g_2$  correlation function

$$\langle \Psi | a_1^{\dagger} a_2^{\dagger} a_1 a_2 | \Psi \rangle + \text{h.c.} = 2 \mathcal{E}_1^2 \mathcal{E}_2^2 N_1 N_2 \text{Re} \left[ e^{i [(\mathbf{k}_2 - \mathbf{k}_1) \cdot (\mathbf{r}_A - \mathbf{r}_B) - (\omega_2 - \omega_1)(t_A - t_B))]} \right]$$
(87)

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Single emitter: antibunching

A single atom emitter and compute

$$G_2(\mathbf{r}, 0, \mathbf{r}, \tau) \equiv G_2(\tau) \tag{88}$$

Use Heisenberg picture,  $E^+$  being proportional to the atomic dipole i.e.  $\sigma_-(\tau)$ 

$$G_2(\tau) = \langle \sigma_+(0)\sigma_+(\tau)\sigma_-(\tau)\sigma_-(0)\rangle \tag{89}$$

Initially  $\sigma_{-}(0) = |g\rangle \langle e|$  and  $\sigma_{+}(0) = |e\rangle \langle g|$ . At time  $\tau$ ,  $\sigma_{\pm}(\tau) = U^{\dagger}\sigma_{\pm}(0)U$  and

$$G_{2} = \left\langle \left| e \right\rangle \left\langle g \right| \left[ U^{\dagger} \left| e \right\rangle \left\langle g \right| U \right] \left[ U^{\dagger} \left| g \right\rangle \left\langle e \right| U \right] \left| g \right\rangle \left\langle e \right| \right\rangle$$
(90)

Evaluate average in |e
angle

$$G_2(\tau) = |\langle e| U(\tau) |g\rangle|^2$$
(91)

hence  $G_2(0) = 0$ 

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A two level atom coupled to a single mode of the radiation field. Coherent coupling larger than dissipative process.

- An atom in an intense laser field
- Cavity quantum electrodynamics

Fruitful to treat atom and mode as a single quantum system. Spontaneous emission and shifts can be added later as a small perturbation.

# The dressed atom model Hamiltonian

$$H = H_a + H'_c + H_{ac} \tag{92}$$

where  $H_a$  and  $H_c' = \hbar \omega_c N$  are the atom and field Hamiltonians. In the RWA

$$H_{ac} = -i\hbar \frac{\Omega_0}{2} \left[ a\sigma_+ - a^{\dagger}\sigma_- \right]$$
(93)

where we introduce the 'vacuum Rabi frequency'  $\Omega_0$  (assumed to be real):

$$\Omega_0 = 2 \frac{d\mathcal{E}_0 \boldsymbol{\epsilon}_d^* \cdot \boldsymbol{\epsilon}_c}{\hbar} \tag{94}$$

Atom-field detuning

$$\Delta_c = \omega_{eg} - \omega_c \tag{95}$$

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# The dressed atom model Uncoupled states



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Eigenenergies and eigenvectors

In the nth doublet

$$H_n = \hbar \omega_c \left( n + 1/2 \right) \mathbb{1} + V_n \tag{96}$$

with:

$$V_n = \frac{\hbar}{2} \begin{pmatrix} \Delta_c & -i\Omega_n \\ i\Omega_n & -\Delta_c \end{pmatrix} = \frac{\hbar}{2} [\Delta_c \sigma_Z + \Omega_n \sigma_Y]$$
(97)

and

$$\Omega_n = \Omega_0 \sqrt{n+1} \tag{98}$$

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Eigenenergies and eigenvectors

**Eigenvalues**:

$$E_n^{\pm} = (n+1/2)\,\hbar\omega_c \pm \frac{\hbar}{2}\sqrt{\Delta_c^2 + \Omega_n^2} \tag{99}$$

with

$$\tan \theta_n = \Omega_n / \Delta_c \tag{100}$$

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Eigenvectors

$$|+,n\rangle = \cos(\theta_n/2) |e,n\rangle + i \sin(\theta_n/2) |g,n+1\rangle |-,n\rangle = \sin(\theta_n/2) |e,n\rangle - i \cos(\theta_n/2) |g,n+1\rangle$$
(101)

The 'dressed states'.

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Dressed states



Resonant case: Rabi oscillation

 $\theta_n = \pi/2$  for all *n* values.

$$|\pm, n\rangle = [|e, n\rangle \pm i |g, n+1\rangle] / \sqrt{2}$$
(102)

Initial state  $|\Psi_e(0)
angle = |e,n
angle$ 

$$|\Psi_e(0)\rangle = [|+,n\rangle + |-,n\rangle]/\sqrt{2}$$
 (103)

Interaction representation with respect to the constant  $\hbar\omega_c(n+1/2)\mathbb{1}$ . At time t

$$\left|\widetilde{\Psi}_{e}(t)\right\rangle = \left[\left|+,n\right\rangle e^{-i\Omega_{n}t/2} + \left|-,n\right\rangle e^{i\Omega_{n}t/2}\right]/\sqrt{2}$$
(104)

In the uncoupled basis

$$\left|\widetilde{\Psi}_{e}(t)\right\rangle = \cos\frac{\Omega_{n}t}{2}\left|e,n\right\rangle + \sin\frac{\Omega_{n}t}{2}\left|g,n+1\right\rangle$$
 (105)

For an atom initially in g

$$\left|\widetilde{\Psi}_{g}(t)\right\rangle = -\sin\frac{\Omega_{n}t}{2}\left|e,n\right\rangle + \cos\frac{\Omega_{n}t}{2}\left|g,n+1\right\rangle$$
 (106)

Non-resonant coupling

Large detuning case (dispersive regime)  $|\Delta_c| \gg \Omega_n$ 



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Non-resonant coupling

$$E_n^{\pm} = (n+1/2) \hbar \omega_c \pm \hbar \left( \frac{\Delta_c}{2} + \frac{\Omega_n^2}{4\Delta_c} \right)$$
(107)

$$\Delta_{e,n} = \hbar(n+1)s_0; \qquad \Delta_{g,n} = -\hbar n s_0 \tag{108}$$

with:

$$s_0 = \frac{\Omega_0^2}{4\Delta_c} \tag{109}$$

Two complementary effects

• Atomic frequency change (light shifts and Lamb shift)

$$\delta\omega_{eg} = (2n+1)s_0 \tag{110}$$

• Atomic state-dependent mode shift (index of refraction effect)

$$\delta_e \omega_c = s_0 \tag{111}$$

$$\delta_g \omega_c = -\delta_e \omega_c = -s_0 \tag{112}$$

#### The dressed atom model Autler Townes splitting

The large field version of the vacuum Rabi splitting.

Atom driven by an intense laser field resonant on the e/g transition Two dressed states with a splitting  $\Omega\sqrt{n}$  nearly independent of photon number *n* for a large coherent field. Dressed states are superpositions of *e* and *g* with equal weights.

Probe the system on the h to g transition where h is a third level.

- For a negligible laser intensity: a single line.
- For a strong laser: two lines corresponding to the excitation of the two dressed levels, separated by the Rabi splitting.

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# The dressed atom model Mollow triplet

Fluorescence of the dressed levels in a strong resonant laser field. Rabi splitting nearly the same for all relevant photon numbers.

Add atomic relaxation by spontaneous emission. Emission possible on all transitions between the levels

A triplet of lines

- Atomic frequency
- Sidebands at the Rabi frequencies